# Gravity and Condensed Matter Physics: Beginnings Of A Dialogue ?

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# Outline

- 1. Some Reminiscences
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- 3. Some Examples Of Connections
- 4. Conclusions

John was my thesis advisor (1985-90) and then my post-doc mentor (1992-94).

The years I spent at Caltech were some of the most formative in my life and have shaped me in very important ways not only as a physicist but also as a human being.

It is therefore an especial privilege and honor for me to participate in these celebrations.



I arrived at a wonderful time at Caltech.

Anomaly cancellation had just been discovered and John Schwarz was all excited. And Caltech has just hired two young professors, John Preskill and Mark Wise.

I soon came under the spell of John and Mark. Both great physicists of course, but also wonderful human beings.

Gentle, and compassionate and always willing to hear the student out.

Some others might not have needed this, but for a student like me, in a strange land, all tongue tied, it made all the difference.

It was only much later, when I had been in the field for some time, that I realized how truly special these qualities are amongsts scientists like us!

John was offering a course, Advanced Quantum Field Theory (Physics 230), that year and I decided to take it. I was keen to impress him in the hope that he would take me on as a student.

Soon a problem surfaced though. John was a great lecturer but the course was extremely demanding and I found myself struggling.

One of the great things about Caltech grad. School was that you could pretty much do whatever you liked. So one by one I dropped all the other courses. But still could not keep up. I was really struggling.

Attempts to ask intelligent question did not meet with much success. John's famous ``W" smile did not help.

This smile was well known to all his students and was endlessly analysed. What did the ``W'' mean – that he approved or disapproved? Had anyone seen the inverse ``W''?.

The smile added to my worry.



Time has not dulled, nor age diminished ...



## On a really good day ...

[Brian Warr had the most detailed critique of the ``W" smile. Alexios Polychronakos once claimed he had seen the ``inverse W". To our question as to whether John was happier or more upset Alexi had only maintained a sphinx like silence, however.]







My worries began to take a toll and I found I had developed a nervous tick in my eye.

Soon I found myself in the clinic of an eye doctor – one Dr. Kislinger.

We sat across a table while he peered into my retina and questioned me. I told him I was a grad. student at Caltech. ``What subject?" Particle Physics" I answered thinking the matter would end there. But it did not.

``Theory or Experiment ?''Kislinger asked.

This seemed like an unusually well informed doctor so I played along.

"What courses are you taking ?" he continued.

``Only one, QFT taught by Professor Preskill", I said.

"What! John is at Caltech now ?" the doctor asked.

By now I was seriously worried that the nervous attack had spread to my brain.

"I know they had Politzer, when did John move from Harvard? Who else do they have?"

We continued in this vein for sometime.

After the exam Kislinger concluded:

``There's nothing the matter with your eyes. I know Preskill well. He teaches a very demanding course. That's your problem. As your doctor, I recommend that you drop it!".

Seeing me dumbfounded the doctor finally decided to enlighten me.

"I was a particle physicist at Chicago but did not get tenure. So became an ophthalmologist.

His confession broke the tension and we both burst out laughing.

Somehow things began to work out and I did get into John's good books at least to the extent of his taking me on as a student.

The years that followed in grad. school were amongsts some of the happiest in my life.

The wonderful poolside parties that John and Roberta hosted helped.



John was not a very hands on advisor.

At the same time he was always happy to discuss physics with his students.

This combination really worked for many of us.

He was immensely popular, with 6-8 students at any time!













John's sense of humor was understated.

It took awhile to catch onto ...

John overheard me the day before my candidacy wondering aloud to a fellow student whether I might be packing my bags at this time the next day to return home.

He walked by, flashed his W smile, and said nothing.

The next day after the exam emerging from the deliberations he was brief ``Those bags. I'd hold off on packing them!".

Most of all, it was John's broad interests, then ranging from Particle Physics to Cosmology and Aspects of String Theory, and his willingness to always discuss and explain that made all the difference.

Lunch in Chandler could easily last two hours. And then one could formally meet John in the office. In fact the discussions would often continue in the corridors, on walks and even in the rest room.

One day John and I were engaged in a discussion about beta functions in the restroom. I think it was about scheme dependence at higher orders in perturbation theory. A rather abstruse topic. At some point in the discussion Gell-Mann entered the rest room and headed for one of the stalls.

Several minutes later when he reappeared we were still there.

Shaking his head Gell- Mann exited saying ``Remember, I was the one who first figured it out and even then it took me less time!".

In 1989 John taught a course on quantum field theory in curved space. Kip also attended. The atmosphere was electrifying.

Many of us learnt about the information puzzle during the discussions.

Some time during the next few months, John came to my office on day and said, ``We should look at extremal Reissner Nordstrom black holes. The information puzzle can be stated very sharply in that context.

The comment was much ahead of its times.

John, with Sidney Coleman and Frank Wilczek, explored the idea of whether additional quantum hair could be a way out of the information puzzle.

We built two-dimensional models and studied the entanglement entropy grow as a black hole evaporates.

Fiola, Preskill, Strominger, SPT

The two dimensional models were introduced by Callan, Giddings, Harvey and Strominger.

Also developed by Russo, Susskind, Thorlacius.

Although, the title of my talk is different. Many of these ideas, Extremal black holes, Hair and entanglement will play an important role in it.

Gravity and Condensed Matter Physics: Beginnings Of A Dialogue

Why should gravity and string theory have anything to do with condensed matter physics ?

String Theory: High energy Condensed matter physics: low energy

Few different ways to explain this.

1) Landscape : String Theory a general framework.

Hopefully Includes the completion of the standard model.

And Much more.



Connection Between Gravity and Cond. Matt. Physics

2) AdS/CFT correspondence

Gravity (String Theory)

In Anti deSitter space

(d+1 dimensions)

Dual or Equivalent

Field Theory

(d dimensions)


#### **Dual Nature of Correspondence**

Makes gravity a useful ``tool" for the study of strong coupled field theories some of which might be relevant in the study of condensed matter physics.

## Dual Nature of Correspondence

Means that intuition acquired from the study of field theories in various branches of physics can be put to use for the study of gravity.

## AdS/CFT Correspondence

A Theory Of Gravity in d+1 Dimensional Anti-deSitter Space Conformal Field Theory in d dimensions

# AdS/CFT

Can be extended so that



## AdS/CFT Correspondence

The field theory can be roughly thought of as living on the boundary of AdS space.

It is a ``Hologram".



•Anti deSitter Space  $AdS_{d+1}$ 

$$ds^{2} = \left(\frac{r^{2}}{R^{2}}\right)\left(-dt^{2} + \sum_{i=1}^{d-1} dx_{i}^{2}\right) + \left(\frac{R^{2}}{r^{2}}\right)(dr^{2})$$

•Characterised by one single parameter: R: Radius of AdS space.

A signal can reach the boundary,  $r \to \infty$  , and bounce back to the interior in finite time  $\Delta t$  .

AdS space is therefore like a finite box.

The boundary theory lives on the boundary of this box.



## Black Hole: Point Like



#### Black Brane: Extended Domain Wall



We will consider two examples in this talk to illustrate how intuition from field theory/Condensed Matter Physics is furthering our understanding of gravity and vice versa.

Example 1)

On Cond. Matt. side: Many phases of nature.

On gravity side: No hair Theorems.

Black Holes/ Branes have no hair because of the intense pull of gravity on the horizon.

These two statements are in ``tension''. The different phases should correspond to different kinds of black branes.

Perhaps something is wrong with the no-hair lore on the gravity side?

Indeed this is what has been found.

Guided by the expectations from field theory many new kinds of black brane solutions have been found.

## New Brane Solutions in Gravity

In the examples here we will mostly discuss zero temperature phases. The corresponding black branes are called Extremal Black Branes.

An interesting phase structure at zero temperature can emerge by varying the chemical potentials for other conserved charges (besides energy).

# Example 1)

Holographic Superconductor (Gubser, Hartnoll, Herzog, Horowitz ...)

Gravity  $\longleftrightarrow$  Field Theory Gauge Symmetry  $\Leftrightarrow$  Global Symmetry Broken Gauge  $\longleftrightarrow$  Broken Global Sym Symmetry?

## Example 1): Holographic superconductor

Natural way to try and break the gauge symmetry is to introduce a charged scalar (the Higgs) in the bulk.

However no hair theorems would suggest that the charged scalar would fall into the black brane and thus not be supported outside.

# Example 1):

The No-Hair Theorems turn out to be not valid in asymptotically AdS space.

In an appropriate range of parameter space one gets new hairy black brane /hole solutions which break the gauge symmetry.

## Black Hole With Charged Scalar Hair



## Holographic Superconductor

This happens in AdS space because it is like a box.

In flat space if one tries to evade the no-hair theorems by increasing the charge of the scalar field, there is dielectric break down of the vacuum near the horizon. Charged scalar particles escape to infinity allowing the black hole to sheds the scalar hair.

#### Holographic Superconductors

But AdS space is like a box.

The charged particles cannot escape and must return.

So that the eventual time independent solution is one with charged scalar hair.

This kind of thinking can be extended to find many new kinds of brane solutions in AdS gravity.

After all there many many different kinds of phases seen in nature.

For example: Brane solutions have been found for which the near-horizon geometry is Homogeneous but anisotropic (breaks rotational invariance).

They correspond to homogeneous phases of matter which break rotational invariance.

lizuka, Kachru, Kundu, Narayan, Sircar, S. P. T, Wang

In fact all such solutions, and corresponding phases of matter, can be classified using the Bianchi classification, first developed in cosmology.

## Example: Spin density wave phase

#### Domokos, Harvey; Nakamura, Ooguri, Park; Donos, Gauntlett



# with: Goldstein, Iizuka, Kachru, Kundu, Narayan, Prakash, Sircar, Westphal

Consider field theory system with an global U(1) symmetry (Think of this as Fermion number).

We will be interested in phases where this symmetry in not broken.

An important property is compressibility:  $\underline{\partial n} \neq$ 

n: charge density  $\mu$ : chemical potential

Essentially only one phase is known in condensed matter physics which is compressible with the global symmetry being unbroken:

Fermi liquid.

(Subir Sachdev)

It is important to explore theoretically if other phases of this type are possible.

Some of these could perhaps help in understanding the High Tc materials for example which are well known to exhibit non-Fermi liquid behaviour.



It turns out that in gravitational systems many such compressible phases can arise quite easily!

Here we will work in 3+1 dimensions in the gravity description.

The field theory will be in 2+1 dimensions.

The simplest example is the extremal Reissner Nordstrom Brane. This is a ``unphysical'' though due to the large entropy at zero temperature.

$$S = \int d^4x \sqrt{-g} [R - \Lambda + \frac{1}{e^2} F^2]$$

#### New Compressible phases

- Interesting but too exotic!
- •Compressible:  $n \propto \mu^2$
- -Entropy density:  $s\propto \mu^2$
- •Possibly unstable.
## **Einstein-Dilaton-Maxwell System**

Instead let us consider a system with an additional neutral scalar, the dilaton.

$$S = \int d^4x \sqrt{-g} [R - \Lambda - e^{2\alpha\phi} F^2 - \frac{1}{2} (\partial\phi)^2]$$

 $\alpha$ : an important parameter

Rocha, Gubser; Goldstein, Kachru, Prakash, SPT

## **Dilatonic Extremal Branes**

Intuition: For fixed charge the stress energy of the Maxwell field can change as the dilaton changes.

If the dilaton ``runs off" to  $\pm\infty$  the gauge coupling could go to zero and the gravitational radius could vanish.

Leading to zero entropy.

## **Dilatonic Extremal Branes**

#### Indeed this turns out to be true.

## **Dilatonic Extremal Brane**

# Compressible: $n \propto \mu^2$

## **Specific heat:** $C_v \sim T^p, p \neq 1$

Not Fermi liquid.

## **Dilatonic Extremal Branes**

This system was generalised by Charmosis, Goteraux, Kim, Kiritsis, Mayer

The fields are the same but the couplings are more general with more parameters.

# Key question: Is there a Fermi surface?

# Key insight: Ogawa, Takyanagi, Ugajin

## Key Insight: Use Entanglement



## Entanglement

Fermi surface:

 $S_{EE} \sim A \log(A)$ 

This same behaviour arises for some range of parameters in the Einstein-Dilaton-Maxwell system.

Specific heat can be understood in terms of gapless excitations which disperse with a dynamic exponent z.

Huijse, Sachdev, Swingle

# **Conclusions: Compressible Phases**

These developments raise the tantalising possibility that some of the compressible phases found in gravity could describe non-Fermi liquids.

Perhaps some of these phases might even be of relevance in nature, although it is too early to tell.

# Conclusions Of The Talk

 It is too early to tell whether a definite and useful connection between the study of gravity and condensed matter physics will be established.

 Some of the initial developments do hold promise and are exciting.

## Conclusions Of The Talk

Such a connection would be a wonderful birthday present for John!

Who has always emphasised that the great ideas of physics are all connected and tied together.

In fact this thought has been very much a part of the ethos at Caltech.

As Feynman put it, ``Nature uses only the longest threads to weave her patterns, so that each small piece of her fabric reveals the organization of the entire tapestry."





## There are string embeddings.

K. Narayan; H. Singh; Narayan, Takaynagi, SPT